

# Utilizing BTDF Window Data

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Anywhere Software

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## What Good Is a BTDF?

- The bidirectional transmittance distribution function (BTDF) describes how light passes through a surface
- Some devices exist for measuring BTDFs
- General ray-tracers can compute BTDFs
- Using BTDFs avoids sampling issues/limits

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# Doesn't Radiance Have a BTDF Material Already?

- ⊗ Yes, but it only works for light sources
- ⊗ Sky contributions are counted as diffuse
- ⊗ This is a poor approximation for many materials
- ⊗ Fully enabling the BTDF type is difficult and would be computationally expensive

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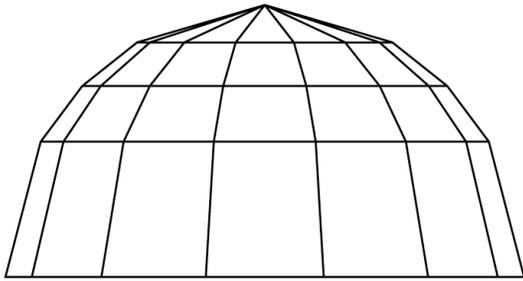
# How to Use a BTDF in Radiance

- ⊗ We can insert the BTDF at the appropriate point in a **mkillum** precalculation
  - ⊗ Special care is required for light sources
- ⊗ Annual calculations require **rtcontrib**

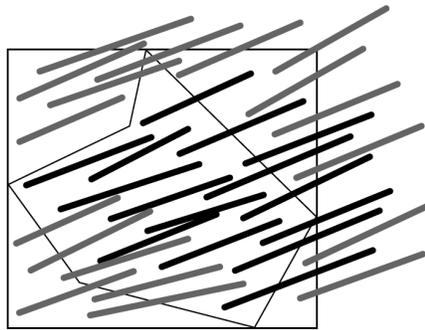
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# Standard mkillum Sampling

Ray samples sent towards illum "front"



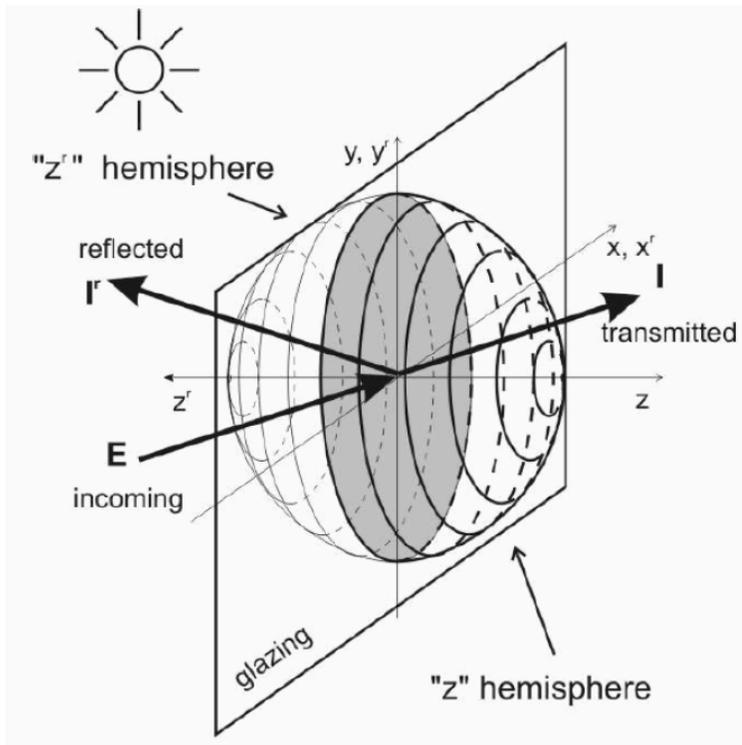
Angle Stratification



Position Distribution

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Klems BTDF angle definitions



145 angles for incident/exiting

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# mkillum BTDF Sampling

- ⦿ Send outgoing rays from opposite side of fenestration according to BTDF angles
- ⦿ Sample rays to light sources if needed
- ⦿ Pass incident light distribution through BTDF
- ⦿ Resample transmitted light and store as window output distribution (as before)

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# New mkillum Settings

- ⦿ Up direction (to orient BTDFs on windows):  
u=[+Z]
- ⦿ Fenestration thickness (for sampling):  
t=[0]
- ⦿ BTDF data file (alt. samples/steradian):  
d=[48]

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# Results

- ☉ Unfortunately, this is a work in progress
- ☉ So far, I've only managed to convert **mkillum** to use library calls rather than **rtrace**
- ☉ Need to add source sampling, read BTDF data

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# Annual Simulations

- ☉ In case you haven't tried it, **mkillum** provides for some rather slow annual calculations
- ☉ Daylight factor method is fab, but **rtcontrib** cannot compute anything **rtrace** cannot
- ☉ BTDFs don't work properly with **rtrace**, so...

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# Three Phase Method

- ⊗ First phase:  
Use **rtcontrib** to get daylight coefficients relating sky patches to incident directions
- ⊗ Second phase:  
Use **rtcontrib** to relate exiting portal directions to \_\_\_\_\_
- ⊗ Third phase (time-step calculation):  
 $\text{sky} * \text{incident} * \text{BTDF} * \text{exiting}$

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# Method Advantages

- ⊗ Precalculation is the only expensive part
- ⊗ Time-step allows fenestration (BTDF) to change or adapt, since matrix multiplication is relatively quick
- ⊗ Anything can be produced, including images

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# Phase 1



**rtcontrib** incoming directions to Tregenza patches

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# Phase 2

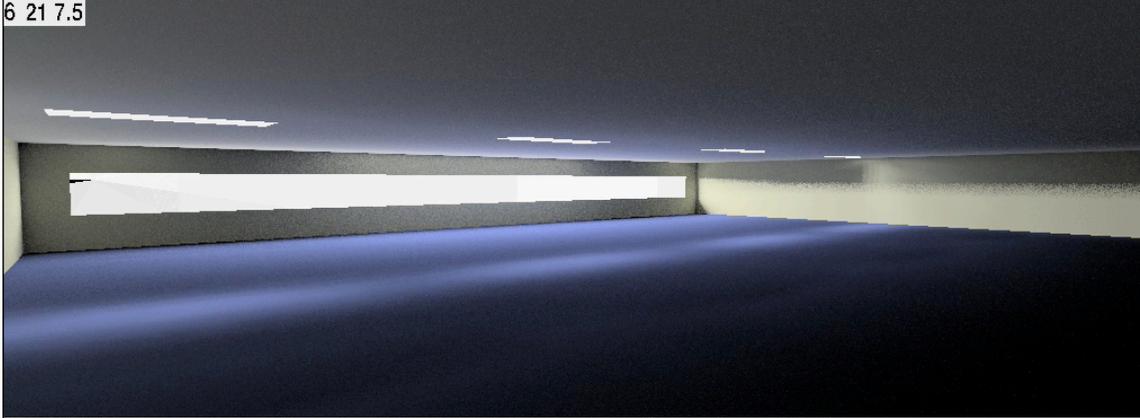


Output side of fenestration to image pixels

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# Phase 3

6 21 7.5



Time-step calculation (matrix multiply)

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## Method Disadvantages

- ⦿ Large-scale external geometry (fins, overhangs, etc.) require subdividing opening
  - ⦿ Best if this can be automated
- ⦿ Limited resolution of BTDF means direct solar will be approximate at best

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# Open Questions

- ⊗ How & when shall we sample light source directions in **mkillum**?
- ⊗ Should we use BRDF as well for interior reflections?
- ⊗ How shall we specify opening subdivision and coordinate a three-phase calculation?
- ⊗ How do we guarantee reliable BTDF input?